

# Position and time of a vessel by celestial observations

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## The observations

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t – time (WT, UTC, UT1, TT)

H – Altitude of a celestial body

LD – Lunar distance

## The equations

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### For an altitude

$H = H(B, L, Dec, GHA) = H(t, B, L)$

$$H = \arcsin(\sin Dec \sin B + \cos Dec \cos B \cos(GHA + L))$$

$$dH = \frac{\partial H}{\partial B} dB + \frac{\partial H}{\partial L} dL + \frac{\partial H}{\partial t} dt$$

$$dH = H_o - H$$

The analytical expressions for altitude derivatives are:

$$\frac{\partial H}{\partial B} = \frac{1}{\cos H} (\sin Dec \cos B - \cos Dec \sin B \cos(GHA + L))$$

$$\frac{\partial H}{\partial L} = \frac{-1}{\cos H} (\cos Dec \cos B \sin(GHA + L))$$

$$\frac{\partial H}{\partial t} = \frac{H_{i+1} - H_i}{t_{i+1} - t_i} = \frac{H(t_{i+1}, B_i, L_i) - H(t_i, B_i, L_i)}{t_{i+1} - t_i}$$

### For a lunar

$LD = LD(t, B, L)$

$$dLD = \frac{\partial LD}{\partial B} dB + \frac{\partial LD}{\partial L} dL + \frac{\partial LD}{\partial t} dt$$

$$dLD = LD_o - LD_c$$

$$\frac{\partial LD}{\partial B} = \frac{LD_{i+1} - LD_i}{B_{i+1} - B_i} = \frac{LD(t_{i+1}, B_i, L_i) - LD(t_i, B_i, L_i)}{B_{i+1} - B_i}$$

$$\frac{\partial LD}{\partial L} = \frac{LD_{i+1} - LD_i}{L_{i+1} - L_i} = \frac{LD(t_{i+1}, B_i, L_i) - LD(t_i, B_i, L_i)}{L_{i+1} - L_i}$$

$$\frac{\partial LD}{\partial t} = \frac{LD_{i+1} - LD_i}{t_{i+1} - t_i} = \frac{LD(t_{i+1}, B_i, L_i) - LD(t_i, B_i, L_i)}{t_{i+1} - t_i}$$

## The solution

In matrix form for three altitudes and a lunar:

$$\begin{bmatrix} dH_1 \\ dH_2 \\ dH_3 \\ dLD \end{bmatrix} = \begin{bmatrix} \frac{\partial H_1}{\partial B} & \frac{\partial H_1}{\partial L} & \frac{\partial H_1}{\partial t} \\ \frac{\partial H_2}{\partial B} & \frac{\partial H_2}{\partial L} & \frac{\partial H_2}{\partial t} \\ \frac{\partial H_3}{\partial B} & \frac{\partial H_3}{\partial L} & \frac{\partial H_3}{\partial t} \\ \frac{\partial LD}{\partial B} & \frac{\partial LD}{\partial L} & \frac{\partial LD}{\partial t} \end{bmatrix} \begin{bmatrix} dB \\ dL \\ dt \end{bmatrix}$$

An initial value for the unknowns is needed: B0, L0, t0

The improved solution is:

$$B = B0 + dB$$

$$L = L0 + dL$$

$$t = t0 + dt$$

Iterate until error < epsilon

### **The motion of the observer**

B – latitude

L – longitude

R – course

V – speed

$$B = B( B_0, L_0, R, V, t-t_0 )$$

$$L = L( B_0, L_0, R, V, t-t_0 )$$

t → Observation

t<sub>0</sub> → Celestial position( B<sub>0</sub>, L<sub>0</sub> ) and Motion( R, V )

$$d = V(t-t_0) = V\Delta t$$

$$B = B_0 + d \cos R = B_0 + V\Delta t \cos R$$

$$L = L_0 + d \frac{\sin R}{\cos B_m} = L_0 + V\Delta t \frac{\sin R}{\cos B_m}$$

$$B_m = (B_0+B)/2 = B_0 + \frac{1}{2}V\Delta t \cos R$$

Changes in R,V for multiple leg tracks: the simplest way to take it into account, is to introduce the adequate equation of restriction for the motion to each observation equation:

$$B = B_0 + \Sigma \Delta B$$

$$L = L_0 + \Sigma \Delta L$$

## Generalization

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If the unknowns are: time, latitude and longitude, the general differential equation of the measured variable F is:

$$dF = \frac{\partial F}{\partial t} dt + \frac{\partial F}{\partial B} dB + \frac{\partial F}{\partial L} dL$$

or in finite differences:

$$\Delta F = \frac{\partial F}{\partial t} \Delta t + \frac{\partial F}{\partial B} \Delta B + \frac{\partial F}{\partial L} \Delta L$$

navigational data from different sources can be combined, F =

1. Altitude
2. Lunar distance
3. Star-star distance
4. Azimuth (measured on land with an astronomical instrument)
5. Bearing
6. Distance
7. Horizontal angle
8. Radio-bearing
9. ...

$$dF = F_{\text{observed}} - F_{\text{calculated}}$$

and the partial derivatives are obtained with finite differences:

$$\frac{\partial F}{\partial x} = (F_{i+1} - F_i) / (x_{i+1} - x_i)$$

$$\begin{bmatrix} 1 \\ 2 \\ dH \\ n \end{bmatrix} = \begin{bmatrix} \frac{\partial F}{\partial B} & \frac{\partial F}{\partial L} & \frac{\partial F}{\partial t} \\ \frac{\partial B}{\partial F} & \frac{\partial L}{\partial F} & \frac{\partial t}{\partial F} \\ \frac{\partial B}{\partial H} & \frac{\partial L}{\partial H} & \frac{\partial t}{\partial H} \\ n \end{bmatrix} \begin{bmatrix} dB \\ dL \\ dt \end{bmatrix}$$

If number of equations  $n > 3$ , the method of the least squares is used to solve the system.