## June 17, 1896 noon sight

Sun's declination at Greenwich mean noon June 17 is given in the 1896 American Ephemerides as N $23^{\circ} 24^{\prime} 50.2^{\prime \prime}$, increasing $3.55^{\prime \prime}$ per hour. Local noon at $138^{\circ} 35^{\prime} \mathrm{W}$ longitude makes GMT some $9.2^{\text {h }}$ (astronomical time) giving an increment of $33^{\prime \prime}$ in declination, let's say $23^{\circ} 25^{\prime}$. Semidiameter of the sun $16^{\prime}$ to the nearest minute of arc.

| $\mathrm{h}_{\text {i }}$ | $55^{\circ} 19^{\prime}$ |  |
| :---: | :---: | :---: |
| i | - 2 |  |
| $\Theta$ | - 3 | (guessed dip value) |
| $\mathrm{ha}_{\mathrm{a}}$ | 5514 |  |
| r | - 1 |  |
| sd | $+\quad 16$ +55 | (LL assumed) |
| H | 5529 |  |
| $\delta$ | $\underline{2325} \mathrm{~N}$ |  |
|  | 8960 |  |
| $\mathrm{H}+\delta$ | 7854 |  |
| $\varphi$ | 116 S | (by DR $11^{\circ} 0^{\prime}$ ) |

## June 17, 1896 pm time sight

The Ephemerides gives Equation of Time $45.61^{s}$ at Gwch mean noon, increasing $0.54^{s}$ per hour. GMT at this time must be around $12^{\text {h }}$, making EoT $52^{\text {s }}$, to be "Subtracted from Mean Time", i.e. added to apparent time. The declination have increased to $23^{\circ} 26^{\prime}$ but in order to keep the sum (called 2 s below) even to avoid interpolation in the log tables we stick to $23^{\circ} 25^{\prime}$. It affects the result by a few seconds of time but compared with the uncertainty in the GMT obtained by the lunar the error is insignificant.

| $\mathrm{h}_{\text {i }}$ | $38^{\circ} 39^{\prime}$ |  |  |
| :---: | :---: | :---: | :---: |
| i | - 2 |  |  |
| $\Theta$ | - 3 |  |  |
| $\mathrm{ha}_{\mathrm{a}}$ | 3834 |  |  |
| r | - 1 |  |  |
| sd | + <br> $+\quad 16$ | (LL assumed) |  |
| h | 3849 |  |  |
| $\varphi$ | 116 | $\log \mathrm{sec}$ | 0.00820 |
| p | 11325 | log csc | 0.03733 |
| 2s | 16280 |  |  |
| s | 8140 | $\log \cos$ | 9.16116 |
| h | 3849 |  |  |
| s-h | 4251 | $\log \sin$ | 9.83256 |
|  |  |  | 19.03925 |
| ½LAT | $1^{\mathrm{h}} 17^{\mathrm{m}} 17^{\text {s }}$ | $\log \sin$ | 9.51963 |
| LAT | 23434 pm |  |  |
| EoT | 52 |  |  |
| LMT | 23526 pm | $\left(17^{\text {th }}\right.$ ) |  |
| GMT | 115048 | (astro $17^{\text {th }}$ ) |  |
| longitude | $9^{\mathrm{h}} 15^{\mathrm{m}} 22^{\mathrm{s}} \mathrm{W}$ |  |  |
|  | $138^{\circ} 51^{\prime}$ | (by DR 138 ${ }^{\circ} 46^{\prime}$ ) |  |

If a table of $\log \sin ^{2}(x / 2)$ with time argument is available the sum of logs (19.03925) can be used to find LAT directly. Then the possible error introduced when halving the $\log$ and doubling the time is eliminated.

