

Define

a	semi-major axis of the orbit
e	orbital eccentricity
s	Sun's radius
ν	the true anomaly
P	the Earth's orbital period

then the semi-minor axis of the orbit is $b = a\sqrt{1-e^2}$

and the Sun-Earth distance is $r = \frac{a(1-e^2)}{1+e\cos\nu}$ which takes extreme values $r_{\max} = a(1+e)$,
 $r_{\min} = a(1-e)$.

From conservation of angular momentum or equivalently Kepler's second law it can be

shown that $\frac{P}{2\pi a^2 \sqrt{1-e^2}} r^2 d\nu = dt$.

In the approximation $\sin^{-1} \frac{s}{r} \approx \frac{s}{r}$ The time averaged semidiameter is

$$\frac{1}{P} \int_0^P \frac{s}{r} dt = \frac{s}{2\pi a^2 \sqrt{1-e^2}} \int_0^{2\pi} r d\nu = \frac{s}{a}.$$

The average of the maximum and minimum semidiameters is $\frac{1}{2} \left(\frac{s}{r_{\max}} + \frac{s}{r_{\min}} \right) = \frac{s}{a(1-e^2)}$.

Proportion of time with $r > a$

Use $r = \frac{a(1-e^2)}{1-e\cos\theta}$ where $\theta = 0$ corresponds to aphelion Then $r = a$ when $\cos\theta = e$. The area

$$\text{swept out while } r \geq a \text{ is } 2 \int_0^{\cos^{-1}e} \frac{r^2}{2} d\theta = -2 \int_1^e \left(\frac{a(1-e^2)}{1-eu} \right)^2 \frac{1}{2\sqrt{1-u^2}} du = a^2 \sqrt{1-e^2} \left(e + \frac{\pi}{2} \right)$$

where the substitution $u = \cos\theta$ has been made. The total area of the ellipse is

$$\pi ab = \pi a^2 \sqrt{1-e^2} \text{ so the proportion is } \frac{1}{2} + \frac{e}{\pi}.$$